

*Errata***Relative figure of merit of optical interferometry and spectroscopy.****Example of parameter estimates of P Cygni circumstellar envelope**M.S. Burgin^{1,2} and A. Chalabaev²¹ Astro-Space Center FIAN, Profsoyuznaya 84/32, 117810 Moscow, Russia

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The on-plot labels indicating the sets of adjustable, nuisance, and target parameters are missing from Figs. 2–6. In what follows, the missing text is specified for each figure. The numbers in parentheses identify the relevant individual plot, the order being from topmost plot to bottom.

Fig. 2. (1) α ; (2) \dot{M} ; (3) v_c ; (4) T_1 .**Fig. 3.** (1) α, T_1 ; (2) α, v_c, T_1 ; (3) $\alpha, v_c, T_1, \Delta_2 T$; (4) $\alpha, v_c, T_1, \Delta_2 T, \Delta_3 T$.**Fig. 4.** (1) α ; (2) v_c ; (3) T_1 ; (4) $\Delta_2 T$; (5) $\Delta_3 T$.**Fig. 5.** (1) α, \dot{M}, v_c ; (2) $T_1, \Delta_2 T, \Delta_3 T$; (3) $\alpha, T_1, \Delta_2 T$; (4) α, \dot{M}, T_1 ; (5) $\alpha, \dot{M}, v_c, T_1, \Delta_2 T, \Delta_3 T$.**Fig. 6.** (1) α ; (2) v_c ; (3) T_1 ; (4) $\Delta_2 T$; (5) $\Delta_3 T$; (6) $\Delta_4 T$; (7) $\Delta_5 T$.**Destriping of polarized data in a CMB mission with a circular scanning strategy**B. Revenu¹, A. Kim¹, R. Ansari², F. Couchot², J. Delabrouille¹, and J. Kaplan¹¹ Physique Corpusculaire et Cosmologie, Collège de France, 11 place Marcelin Berthelot, 75231 Paris Cedex 05, France² Laboratoire de l'Accélérateur Linéaire, IN2P3 CNRS, Université Paris Sud, 91405 Orsay, France*Astron. Astrophys. Suppl. Ser.* **142** (2000) 499-509

New caption of Fig. 8: The Mollweide projection of the residuals of the I –Stokes parameter after destriping of $1/f$ noise plus white noise with $\eta = 1$

New caption of Fig. 9: The Mollweide projection of the residuals of the I –Stokes parameter after zeroing the average of the circles, for $1/f$ noise plus white noise with $\eta = 1$

New caption of Fig. 10: The Mollweide projection of the residuals of the Q –Stokes parameter for a white noise mission